

national accelerator laboratory

EXP-23 September 12, 1972

ACCELERATOR EXPERIMENT -- Radial Dependence of Tune in the Main Ring at 150 GeV

Experimentalists: R. Stiening

Date Performed: 3 & 5 September 1972

The data are appended to this report, followed by Stiening's plot of the tunes versus $\delta p/p$. The last page is a v_v versus v_v presentation of the data, with third order resonance lines sketched in.

We now have two high energy measurements of radial tune dependence. The first, at 74 GeV, was reported in EXP-22. initial comments about the two sets of data are in order; a more detailed discussion will appear in a later report.

First, let's look at v_x versus $\delta p/p$. In EXP-22, we used an "eyeball" measure of the curvature to obtain an estimate of the octupole moment of the quadrupoles. Now, S. Ohnuma has done a least squares fit to second order in $\delta p/p$ for both the earlier data and the present measurements; he finds

$$v_{x} = 20.28 - 0.21 \left(\frac{\delta p}{p}\right)_{s} + 0.15 \left(\frac{\delta p}{p}\right)_{s}^{2}$$
 at 75 GeV

$$v_{x} = 20.29 - 0.21 \left(\frac{\delta p}{p}\right)_{g} + 0.23 \left(\frac{\delta p}{p}\right)_{g}^{2}$$
 at 150 GeV

His result at 75 GeV indicates a larger octupole moment than that deduced in EXP-22, though within the limits of uncertainty of that report. The 150 GeV fit implies an octupole moment still larger by a factor of more than 1.5. As designed, the main ring quadrupoles should not exhibit an increase of positive octupole moment

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with excitation. A more likely source of the difference is a radial dependence of the curvature. A comparison of the RF system frequencies for $\delta p/p=0$ in the two sets of data indicates the lower energy data explored a region shifted to smaller radius by about 0.1% in $\delta p/p$. Thus the curvature seems to increase with increasing radius. Indeed, the plot at 150 GeV does appear to become more linear at smaller momentum. We conclude that it is an over-simplification to try to account for the high energy radial tune dependence by solely the chromatic aberration of the quadrupoles and their octupole moment.

Second, consider ν_y versus $\delta p/p$. Ohnuma's fit to the 150 GeV data is

$$v_{y} = 20.18 - 0.26 \left(\frac{\delta p}{p}\right)_{g} - 0.036 \left(\frac{\delta p}{p}\right)_{g}^{2}$$
 150 GeV

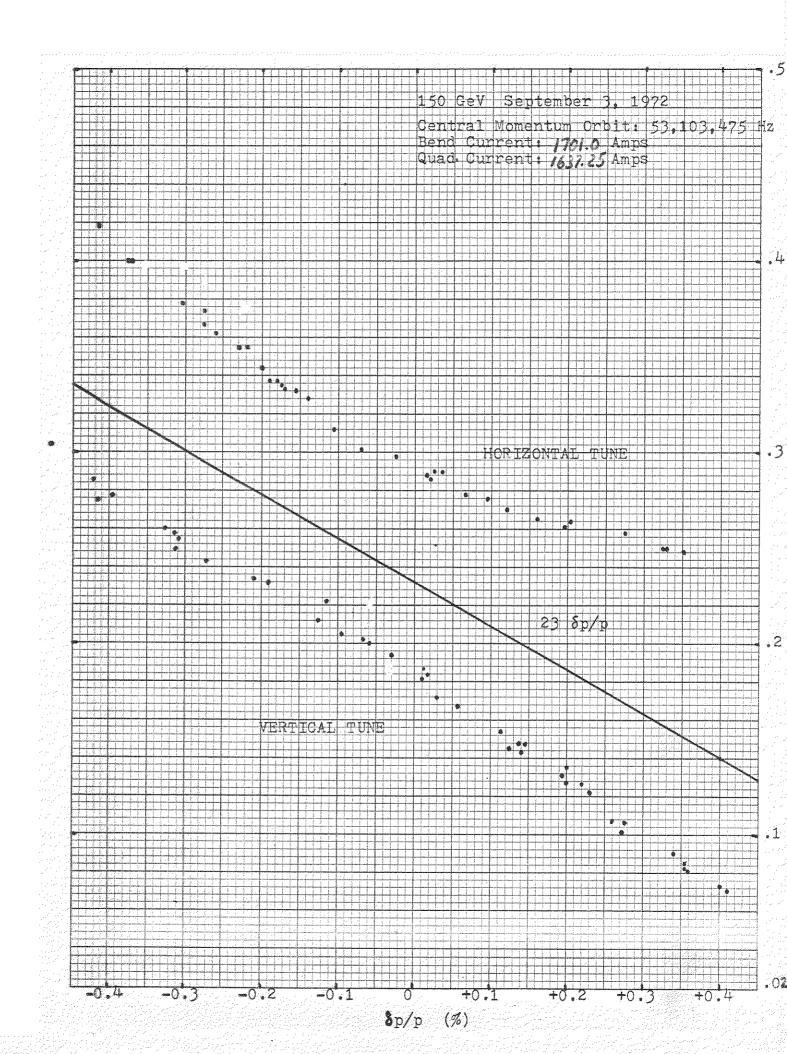
As calculated in EXP-22, the second order term in $\delta p/p$ for vertical motion should be only some 4% of the corresponding term for radial motion if those terms arise from the octupole moment of the quadrupoles. The coefficient above, -.036, is too large by a factor of four and is of the wrong sign. Again, we are forced to conclude that our model is too simple. In fact, the vertical plot looks rather like two straight lines, intersecting at $\delta p/p = 0$.

Third, the v_y versus v_x plot exhibits a discontinuity in slope at $\delta p/p = 0$. Again, we have the intersection of two straight lines, whereas chromatic aberration and an octupole moment common to all the quads would result in a parabola.

Finally, though the trajectory on the ν_y versus ν_x plot crosses a number of third order resonance lines, only at $2\nu_y$ - ν_x = 20

was there an observable effect -- in particular the vertical oscillation excited by the pinger was seen to damp in 60 turns in contrast to continuing for a long time. This coupling resonance would arise from a twentieth harmonic of a sextupole moment, and given an octupole moment in the quads, the twentieth harmonic of the closed orbit distortion (which we know to exist) will produce a twentieth sextupole harmonic.

D. A. Edwards



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